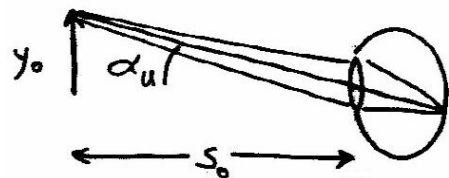


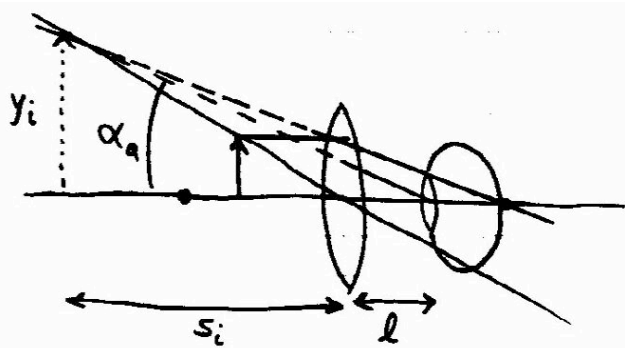
Magnifying Glass (simple magnifier)

unaided eye: image formed on retina.



- get clear image only for $s_o > d_o$
- re. object at near point ($d_o \approx 25 \text{ cm}$)

positive lens: form virtual, erect image ($s_o < f$), with $|s_i| > d_o$:



- rays appear to come from virtual image, and are focussed by eye.

- compare angles subtended by "unaided" image, α_u , with that of "aided" image, α_a ;

(paraxial) $\alpha_u \approx \frac{y_o}{d_o}$, $\alpha_a \approx \frac{y_i}{|s_i| + l}$
 (max value, $s_o = d_o$)

magnifying power $MP \equiv \frac{\alpha_a}{\alpha_u} = \frac{y_i}{y_o} \frac{d_o}{|s_i| + l}$

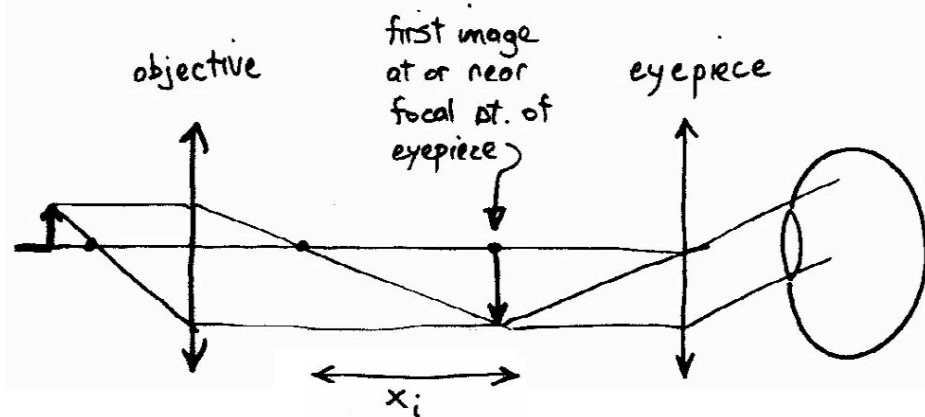
but $\frac{y_i}{y_o} = -\frac{s_i}{s_o} \Rightarrow MP = -\frac{s_i}{s_o} \frac{d_o}{l - s_i}$

get various MP depending on s_i and l ;
 max MP for $l = 0$, and take $|s_i| \rightarrow \infty$
 (for relaxed eye) yielding $s_o = f \Rightarrow$

$$MP \rightarrow \frac{d_o}{s_o} = \frac{d_o}{f}$$

eg. $f = 10 \text{ cm}$, $MP = \frac{25 \text{ cm}}{10 \text{ cm}} = 2.5 \times$
 larger MP limited by aberrations.

Microscope - 2 lenses $f_o \ll f_e$



$$M_{T_o} = -\frac{x_i}{f_o} \quad x_i \equiv \text{tube length, usually standard value of 16 cm}$$

$$\Rightarrow \text{total } M_T = M_{T_o} \cdot M_{P_e} = -\frac{16}{f_o} \cdot \frac{25}{f_e}$$

eyepiece: compound lens - designed for comfortable eye location. MP=10 typical

objective: eg. $f_o = 3.2 \Rightarrow 5\times$ mag.

$$\text{fixed } x_i: s_o = \frac{f_o s_i}{s_i - f_o} = \frac{f_o (x_i + f_o)}{x_i} = f_o + \frac{f_o^2}{x_i}$$

$\Rightarrow s_o \approx f_o$; higher mag \Rightarrow smaller f_o and s_o

Telescopes - A) Astronomical

- focusses rays from object $s_o \rightarrow \infty$

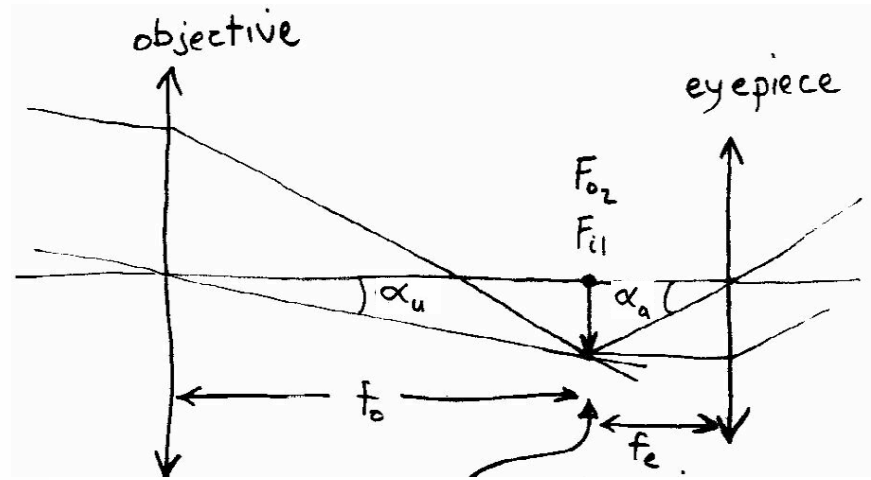


image from objective at F_{i1} ; coincides nearly with F_{o2}

inverted image

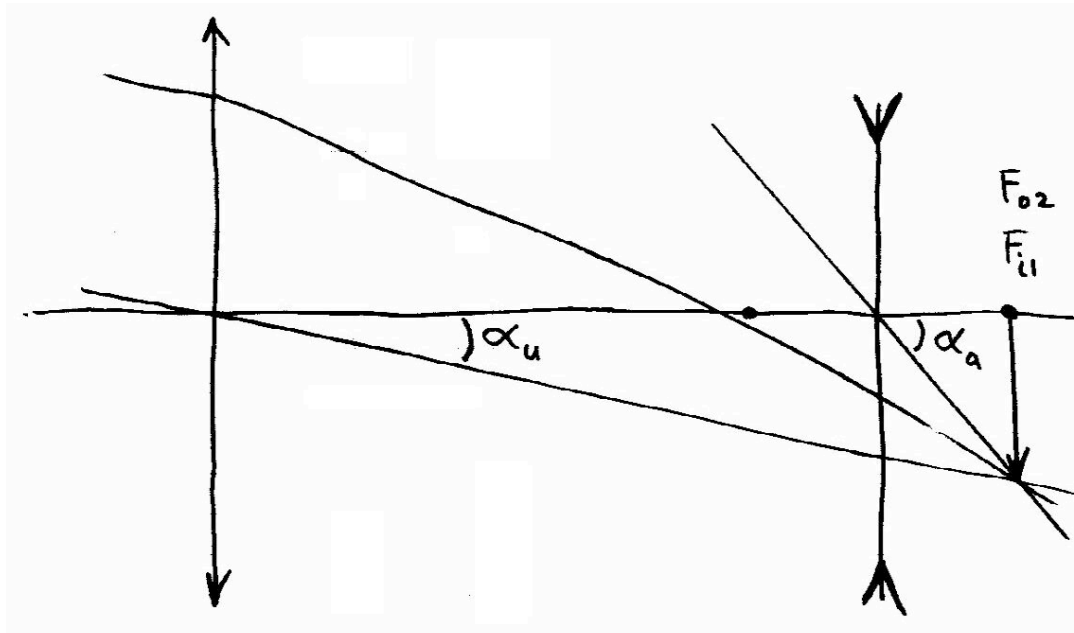
$$M \equiv \frac{\alpha_a}{\alpha_u} = -\frac{f_o}{f_e}$$

$$f_o \gg f_e$$

B) Galilean Telescope

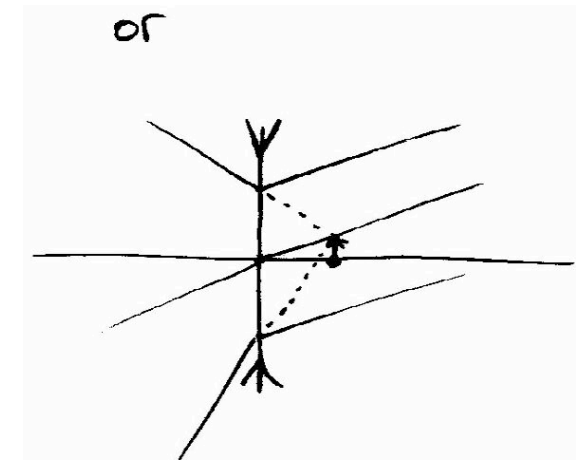
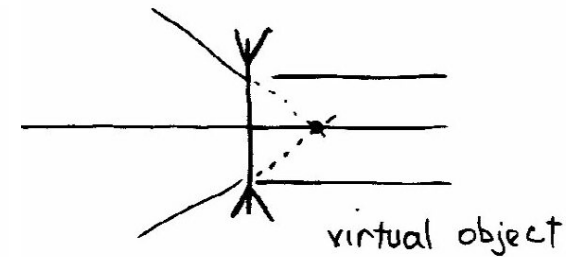
$f_e < 0 \Rightarrow$ erect image

length = $f_o + f_e$, shorter than for astronomical telescope



again, $M \equiv \frac{\alpha_a}{\alpha_u} = -\frac{f_o}{f_e}$

recall: negative lens:

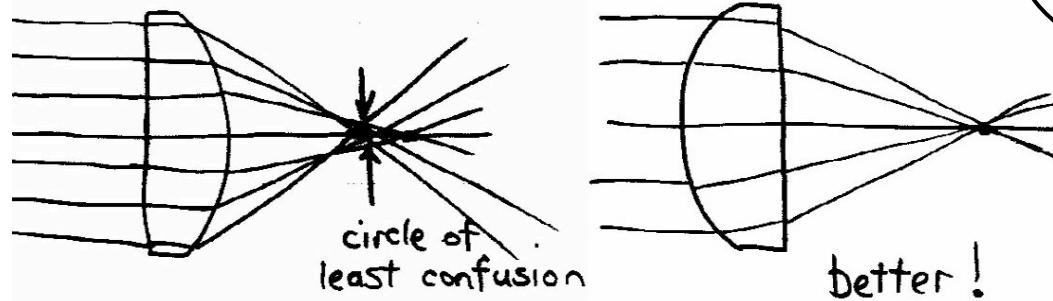


Aberrations: arise from -

- (a) paraxial approx $\left\{ \begin{array}{l} \text{spherical aber.} \\ \text{coma} \\ \text{astigmatism} \\ \text{curvature of field} \\ \dots \end{array} \right.$
- (b) dep. of n on λ

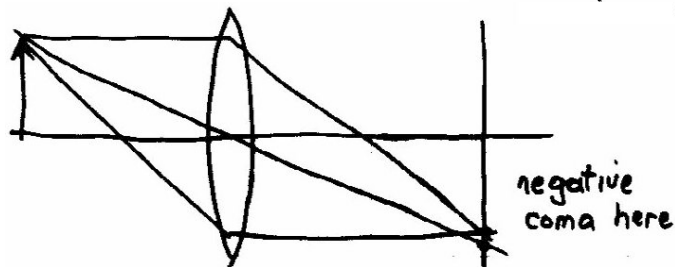
examples:

- spherical aberration: nonparaxial rays are too strongly bent

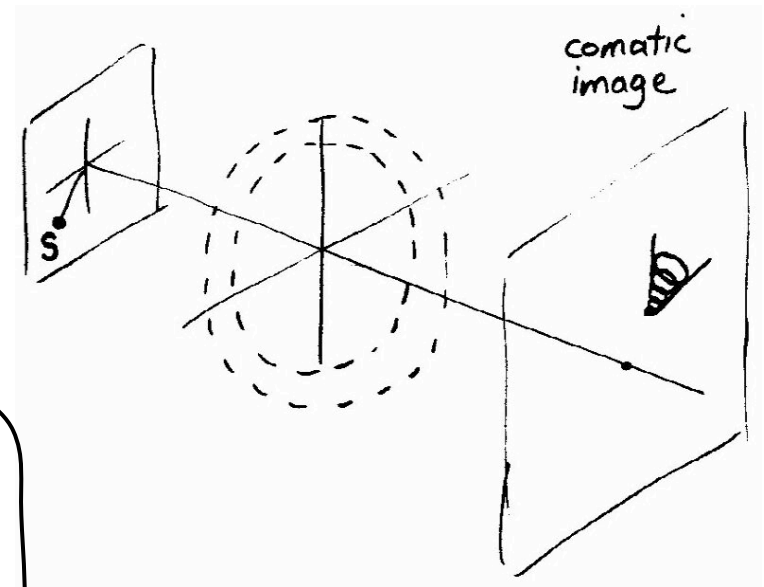


- coma: arises when $y_0 \neq 0$

also depends on lens shape.

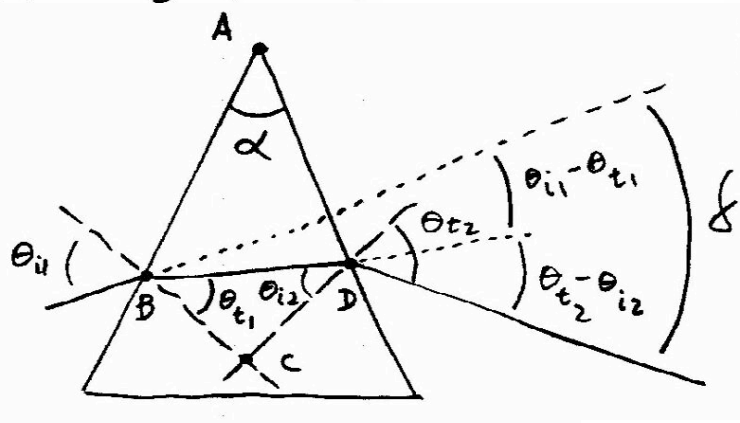


consider focussing of off-axis point:



Prisms - variety of types

dispersing prism:



given α , n , θ_{i1} determine θ_{t2} :

look at total deviation δ :

$$\delta = (\theta_{i1} - \theta_{t1}) + (\theta_{t2} - \theta_{i2})$$

- polygon ABCD has 2 right angles

$$\Rightarrow \alpha + \angle BCD = \pi$$

$$\text{and } \angle BCD = \pi - (\theta_{t1} + \theta_{t2})$$

$$\Rightarrow \alpha = \theta_{t1} + \theta_{t2}$$

$$\text{thus } \delta = \theta_{i1} + \theta_{t2} - \alpha \quad (*)$$

get θ_{t2} from Snell's law:

$$\theta_{t2} = \sin^{-1}(n \sin \theta_{i2}) = \sin^{-1}(n \sin(\alpha - \theta_{t1}))$$

$$\frac{\sin \alpha \cos \theta_{t1} - \cos \alpha \sin \theta_{t1}}{(1 - \sin^2 \theta_{t1})^{1/2}}$$

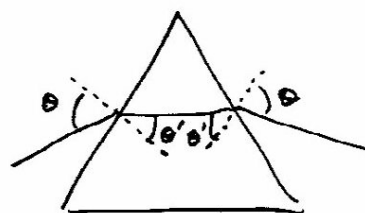
$$\text{and } n \sin \theta_{t1} = \sin \theta_{i1}$$

$$\Rightarrow \theta_{t2} = \sin^{-1} \left[\sin \alpha (n^2 - \sin^2 \theta_{i1})^{1/2} - \cos \alpha \sin \theta_{i1} \right]$$

and get δ from $(*)$.

angle of minimum deviation:

symmetric ray:



$$\theta_{t1} = \theta_{i2} = \theta'$$

$$\theta_{t2} = \theta_{i1} = \theta$$

$$\text{then } \delta = 2\theta - 2\theta'$$

$$\sin \theta = n \sin \theta'$$

$$\alpha = 2\theta' \Rightarrow \theta' = \frac{\alpha}{2}$$

$$\theta = \frac{\delta + 2\theta'}{2} = \frac{\delta + \alpha}{2}$$

can solve for δ (given n)
or n (given δ).

$$\Rightarrow \sin\left(\frac{\delta + \alpha}{2}\right) = n \sin\left(\frac{\alpha}{2}\right)$$